

## Chapter 6 Telescopes: Portals of Discovery



### What are the most important properties of a telescope?

1. **Light-collecting area:** Telescopes with a larger collecting area can gather a greater amount of light in a shorter time.
2. **Angular resolution:** Telescopes that are larger are capable of taking images with greater detail.
3. **Magnification:** How much larger the image appears in the telescope versus the sky.

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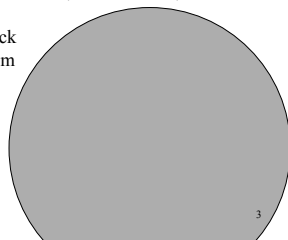
### Light Collecting Area

- A telescope's diameter tells us its light-collecting area:  $\text{Area} = \pi (\text{diameter}/2)^2$

HLCO  
1 m

Keck  
10 m

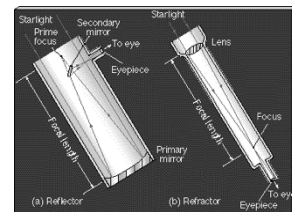
Your eye →  
7 mm



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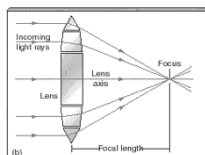
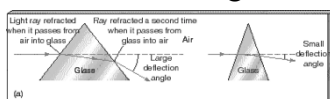
### What are the two basic designs of telescopes?

- **Refracting telescope:** Focus light with lenses
- **Reflecting telescope:** Focus light with mirrors



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### Refracting Telescope



- Refract (bend) light like prisms
- Focal length = distance from lens to focus
- Add eyepiece lens to magnify image

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### Refracting Telescope



- Focal ratio = focal length / lens diameter
- Magnification =  $\text{FL}(\text{telescope}) / \text{FL}(\text{eyepiece})$
- Example:  
 $\text{FL}(\text{telescope}) = 100 \text{ cm}$  and  
 $\text{FL}(\text{eyepiece}) = 2 \text{ cm}$ .  
 $\text{Mag} = 100/2 = 50\times$
- Binoculars:  
"7x50" means  $\text{Mag} = 7\times$ ,  
lens is 50 mm in diameter

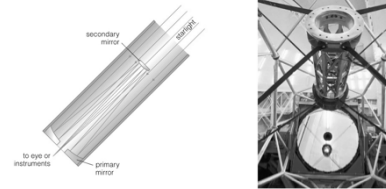
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## Refracting Telescope Problems

- Glass heavy – sags!
- Glass absorbs some colors
- Two surfaces to polish
- Chromatic aberration – different colors focus at different distances
- Better solution: use mirrors!

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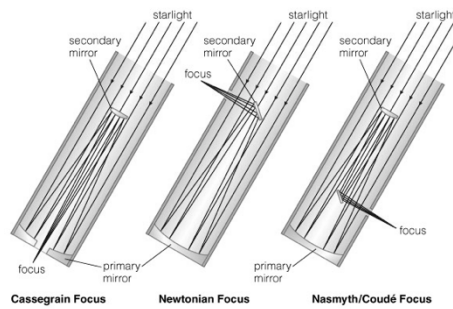
## Reflecting Telescope



- Only one surface to grind and polish
- Support mirror in back so can make larger
- No chromatic aberration
- Need second mirror to reflect light out of tube

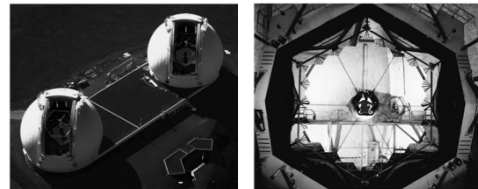
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## Designs for Reflecting Telescopes



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## Mirrors in Reflecting Telescopes

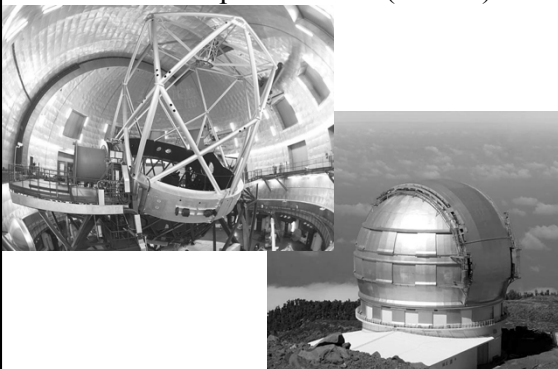


Twin Keck telescopes on Mauna Kea in Hawaii

Segmented 10-meter mirror of a Keck telescope

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## Gran Telescopio Canarias (10.4 m)



## Very Large Telescope (4x8 m)



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## How does Earth's atmosphere affect ground-based observations?

- The best ground-based sites for astronomical observing are
  - Calm (not too windy)
  - High (less atmosphere to see through)
  - Dark (far from city lights)
  - Dry (few cloudy nights)
- Often atop remote mountains

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## Calm, High, Dark, Dry



Summit of Mauna Kea, Hawaii

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## Light Pollution

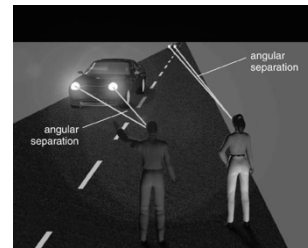


- Scattering of human-made light in the atmosphere is a growing problem for astronomy

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## Angular Resolution

- The *minimum* angular separation that the telescope can distinguish (amount of detail)
- Smaller (better) with larger telescope diameter and smaller wavelength of light



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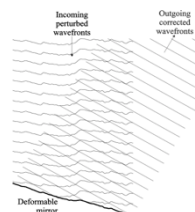
## Atmospheric Turbulence



- Stars appear as twinkling dots
- Blur size typically about 1 arcsecond

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## Adaptive Optics



OFF

ON

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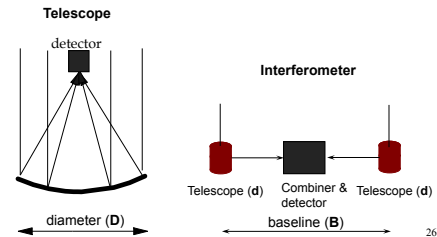
## Center for High Angular Resolution Astronomy (CHARA)

- Long term program at GSU to study tiny details in the sky (“high resolution”)
- Limited by diameter of telescope
- GSU developed novel approach using multiple-telescopes:  
**Long baseline interferometry**
- CHARA Array: best of its kind!

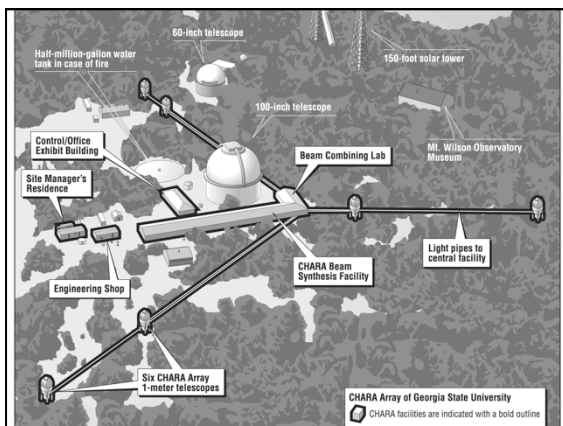
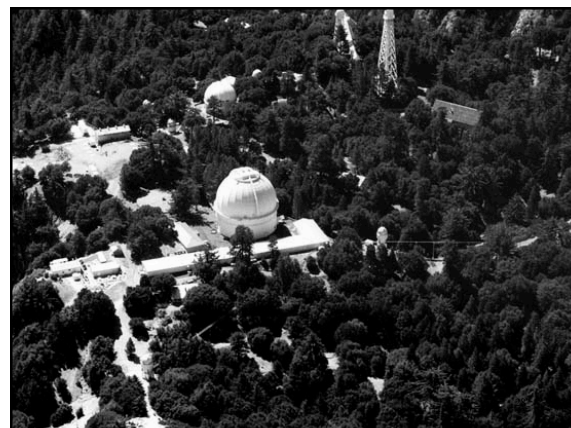
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## Interferometry

- An interferometer combines the light from several small telescopes to yield the power of a much larger telescope
- The larger the separation, the better the results



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## Six Telescopes, Largest Baseline of 330 m

### Telescope Enclosures

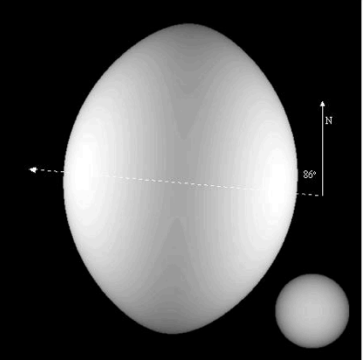


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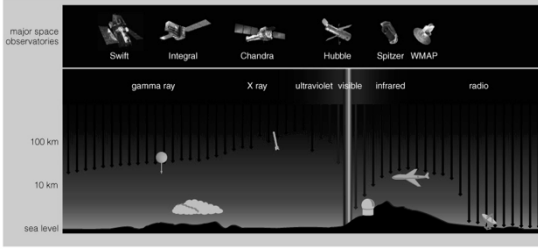
**Resolve objects the size of a dime in LA as viewed from Atlanta**

**Rapidly spinning star Regulus**




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### Transmission in Atmosphere



- Only radio and visible light pass easily through Earth's atmosphere
- We need telescopes in space to observe other forms


### Radio Telescopes



- A radio telescope is like a giant mirror that reflects radio waves to a focus
- Arecibo, Puerto Rico; 205 m dish

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### Radio Telescopes



- Observe interstellar clouds invisible in optical
- Interferometry easier in radio; Very Large Array (VLA), New Mexico

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### Radio Telescopes



- Atacama Large Millimeter Array 64 radiotelescope interferometer in the foothills of Andes Mountains in Chile
- Detailed images of cool gas and dust in universe

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### Infrared Telescopes

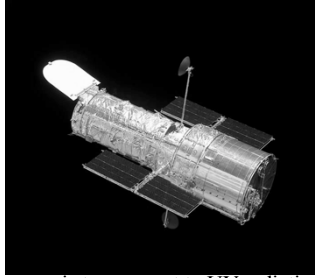


SOFIA      Spitzer

- Infrared telescopes operate like visible-light telescopes but need to be above atmosphere
- View cool, dusty objects in space

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## Ultraviolet: Hubble Space Telescope



- Outer space is transparent to UV radiation
- Sky is very dark, no atmospheric turbulence
- Observe very hot and faint objects

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## X-Ray Telescopes

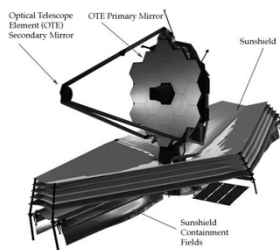


Chandra

- X-ray telescopes also need to be above the atmosphere
- Observe very hot gas (millions of degrees)

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## Future Telescope



- James Webb Space Telescope is a large (6.5 m), infrared-optimized space telescope, scheduled for launch in 2018
- Will observe redshifted light from distant galaxies

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## Next time:

- Chapter 7:  
Solar System  
please read pages 191 – 209 in text.